which is large in the *D*-region, has been neglected for the purpose of calculation. It is possible that a full-wave analysis or a short-interval ray tracing approach would give significantly different results.

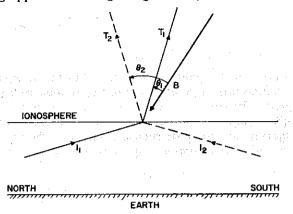


Fig. 2. Transmitted wave normals in the magnetic meridian.

Another explanation might be that irregularities cause some scattering that reduces the effective difference in the wave-normal directions. In spite of these uncertainties, it is clear that the direction of propagation with respect to the earth's magnetic field has a marked effect on the intensity of signals which emerge from the absorbing region.

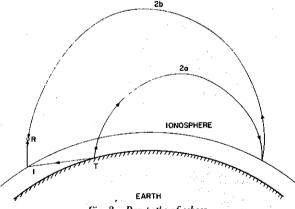


Fig. 3. Ray paths of echoes.

GROUP DELAYS

By noting the delays of the signals, it was usually possible to deduce the general nature of the path of propagation, and to give a picture of the ionosphere consistent with other observations. The "direct" signals (least delay, such as path 1 in Fig. 3) in the northern hemisphere are the

simplest signals to interpret. A typical mid-day delay is 85 m at a height of 600 km. If we assume a Chapman distribution with a maximum plasma frequency of 12 Mc/s at height of 300 km, we obtain a scale height of roughly 120 km. This value is not inconsistent with the average value of 100 km estimated from a number of rocket measurements.⁵

The southern hemisphere group delay measurements show both a direct signal and a "one-hop" whistler mode signal (such as path 2a in Fig. 3). Figure 4 shows the delay data as a function of local mean time. Also shown are the average values of f_0F2 at the equator. The

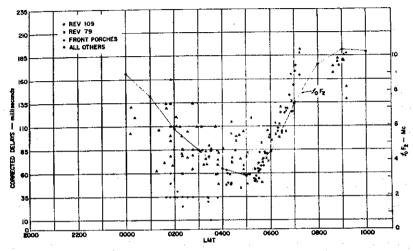


Fig. 4. Group delay and foF2 vs. LMT.

correlation is reasonably good especially in the sunrise period. The data indicate that the scale height increases with time from 80 km at 5.30 a.m. to 100 km at 7 a.m. An interesting result of the analysis is that the exact ray path need not be known to determine the scale height. If the ray path is longer it must reach a higher altitude but the reduced density at the higher altitudes gives less delay per unit path length. These factors almost completely cancel one another at the altitudes involved.

Measurements of the Doppler frequency shift of these signals show a scatter of a few cycles per second with zero mean until about local sunrise. At this time the average frequency suddenly decreases by as much as 7 c/s, where it remains until signals are lost. The observed frequency shift is consistent with the value predicted from the rate of change of group delay using the quasi-longitudinal approximation of the magnetoionic theory.

group by a first factor spring that it is be sur-

Strong two-hop whistler mode signals were frequently observed in the northern hemisphere (Fig. 5). The two ray paths which are postulated to explain the signals are shown in Fig. 3. From Figs. 2 and 3 it is seen

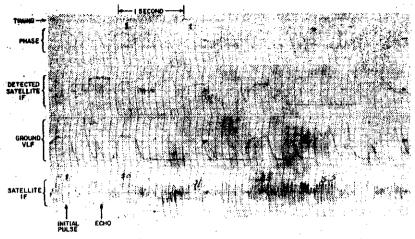


Fig. 5. Echoes on revolution 513.

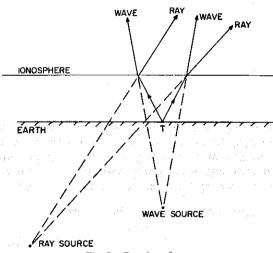


Fig. 6. Beaming of rays.

that the direct signal penetrates the ionosphere with an unfavourable wave-normal angle (see the previous section), and therefore suffers relatively more attenuation. To explain the relatively high strength of the echo, it is necessary to postulate that the first segment of the two-hop path begins close by but to the south of the transmitter. Furthermore, as indicated in Fig. 6, the high refractive index and the strong

garth and agreem states at against come agents

anisotropy of the ionosphere in the whistler mode combine to restrict markedly the divergence of the rays. Using a reasonable model of the ionosphere, the calculated group delay is in good agreement with that measured.

CONCLUSIONS

- (1) The LOFTI-I data clearly show that ducts of enhanced ionization are not required for the propagation of whistler-mode signals to a satellite.
- (2) The differential attenuation measurement demonstrates the pronounced effect of the tilt of the earth's magnetic field on attenuation rate. The results are qualitatively in agreement with the predictions of magnetoionic theory. Note that little or no effect would be expected with either a vertical or a horizontal magnetic field.
- (3) Two-hop whistler mode propagation to a satellite yields signals that are often comparable in amplitude to the direct signal. The echo strength may even, on occasion, exceed the direct signal in amplitude.
- (4) Analysis of the group delay data shows that a scale factor on electron distribution can be obtained from this type of experiment.

REFERENCES

- ¹ LEIPHART, J. P., ZEEK, R. W., BEARCE, L. S., and TOTH, E., Penetration of the ionosphere by very-low-frequency radio signals—interim results of the LOFTI I experiment, *Proc. Inst. Radio Engrs.*, N.Y. 50, 6-17, 1962.
- ² Storey, L. R. O., An investigation of whistling atmospherics, *Phil. Trans. A*, **246**, 113–141, 1953.
- ³ YABROFF, I. W., Computation of whistler ray paths, *J. Res. Nat. Bur. Stand.* **65D**, 5, 485-505 1961.
- ⁴ Budden, K. G., Radio Waves in Ionosphere, Cambridge University Press, 1961.
- WRIGHT, J. W., A model of the F-region above h_{max} F2, J. Geophys. Res. 65, 1, 185-191, 1960.

CHAPTER 7

AN INTERPRETATION OF LOFTI-I V.L.F. OBSERVATIONS*

Louis H. Rorden,† R. L. Smith and R. A. Helliwell;

During February and March of 1961, the Naval Research Laboratory LOFTI-I satellite received very low frequency signals transmitted into the ionosphere from U.S. Navy stations NBA in the Canal Zone and NPG in Jim Creek, Washington. The signals were sometimes observed to have been delayed by as much as 600 milliseconds.

It was possible to deduce the thickness (scale height) of the ionosphere above the height of maximum electron density. The values correspond roughly to those previously found from rocket measurements.

Simultaneous daytime reception of both NPG and NBA showed an interesting difference in attenuation of the signal strengths such that the signal from NBA was more heavily attenuated. The difference can be explained by taking into account the tilt of the earth's magnetic field and the arrival of the NBA signal from the south and the NPG signal from the north.

In general, the whistler mode signals observed in LOFTI-I show no evidence of having been trapped in columns of enhanced ionization. This is in contrast to the characteristics of naturally occurring whistlers observed on the ground, which are best explained by assuming that they have been trapped in such columns. The difference may be explained by noting that whistler energy not trapped in columns of enhanced ionization is not observed on the ground, presumably as a result of total internal reflection at the lower boundary of the ionosphere.

INTRODUCTION

During February and March 1961, the Naval Research Laboratory satellite LOFTI-I received very low frequency (v.l.f.) signals transmitted into the ionosphere from U.S. Navy stations NBA in the Canal Zone and NPG in Jim Creek, Washington.¹ These signals were received in the satellite with a small loop antenna and narrow band receiver (25 cycle bandwidth). Signals were strong at night and generally weak during the daytime, but were observed over every region of the satellite orbit within the range of telemetry recording stations.

† Stanford Research Institute.

The state that the second responsible to the second second second second second second second second second se

Consider the second second second

and the control of the first particles are the property of the control of the con

in the sign of the control of the art here is a first of the

^{*} This work was supported by the Naval Research Laboratory under Contract Nonr 3267(00) (X).

Radioscience Laboratory, Stanford University.

explained. Attenuation calculations were made for a slab model of the is increased, and hence the difference in attenuation is qualitatively markedly as the angle between the wave normal and the earth's field It is to be expected that the attenuation of v.l.f. waves will increase Figure 2 shows the general nature of the wave-normals for the two cases. from north to south for medium latitudes in the northern hemisphere. will be much larger for signals propagating from south to north than The angle of wave-normal refraction with respect to the magnetic field the D-region, where the refractive index is relatively low (about 2.5).

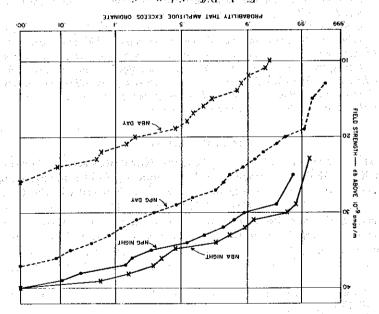


Fig. 1. Differential attenuation.

differential absorption is not actually realized. total absorption—is closer to 25 km, it appears that the calculated above. Since the effective thickness of the absorbing region-based on of 7 km, assuming a constant value of differential absorption as computed is about 10 dB, which corresponds to an effective absorbing path length 1-6 dB/km for MPG and 3-0 dB/km for MBA. The measured difference geneous wave structure.4 The results give an attenuation rate of or magnetic field = 45 angles), including the effects of the inhomolower ionosphere $(f_0 = 285 \text{ kc/s}, f_H = 100 \text{ kc/s}, v = 2 \times 10^6/\text{sec}, \text{ filt}$

region. In addition, the variation of refractive index with altitude, in error because of inaccurate knowledge of the parameters of the Dfractive index at the altitude of maximum absorption, which could be One possible source of this discrepancy is the assumed value of re-

> occasions, the average Doppler shift was as high as 7 c/s. shifts of a few cycles per second were observed consistently. On some observed in a region centred to the north of NBA. Random Doppler 600 msec. The longer delays were attributed to echoes, which were show group delays ranging from a few milliseconds to as high as from the transmitter. Signals were observed to fade randomly and to NBA signals were observed out to a ground range of over 16,000 km

PATHS OF PROPAGATION

components, within an hour or so, and the integrally related time intervals of echoing the multipath character of whistlers, the similarity of different whistlers earth's magnetic field. This hypothesis has been advanced to explain the signals are trapped in ducts of enhanced ionization aligned with the Analysis of ground-based whistler observations strongly suggests that

undetected. wave-normal direction and field intensity might well have gone the satellite had been in a whistler duct, any characteristic changes in pretation of the data was made without requiring the ducts. Even if paths not aligned with the earth's magnetic field. A consistent interpropagation—under certain conditions at least—takes place along propagation was found. On the other hand, evidence was found that ducts (estimated from the ground-based results). No evidence of ducted over distances greater than the average spacing between the postulated satellite. These signals, moreover, were often observed continuously Very low frequency signals were observed over a wide area at the

Storey,2 and by Yabroff,3 guided to a large extent by the earth's magnetic field as discussed by reflection at the boundary. The non-ducted mode is nevertheless the ionosphere, whereas the energy not so trapped suffers total internal trapped in ducts is readily transmitted across the lower boundary of characteristics and the LOFTI results appears to be that whistler energy One explanation for the difference in the interpretation of whistler

DIFFERENTIAL ATTENUATION

signals were about the same at the satellite. time conditions, on the other hand, the intensities of NBA and NPG strengths from these two stations were about the same. Under nightthought the satellite was near a ground receiver at which the field from NPG were consistently stronger than those from NBA, even for both day and night are shown in Fig. 1. In the daytime, the signals east of NPG. The cumulative amplitude distributions of the two signals stations NPG and NBA in a region generally north of NBA and south-An interesting result was the simultaneous observation of signals from

The greatest attentuation of the v.l.f. signals in the daytime occurs in

which is large in the *D*-region, has been neglected for the purpose of calculation. It is possible that a full-wave analysis or a short-interval ray tracing approach would give significantly different results.

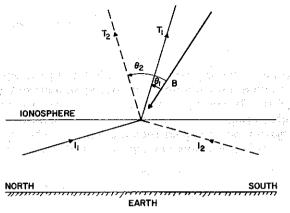


Fig. 2. Transmitted wave normals in the magnetic meridian.

Another explanation might be that irregularities cause some scattering that reduces the effective difference in the wave-normal directions. In spite of these uncertainties, it is clear that the direction of propagation with respect to the earth's magnetic field has a marked effect on the intensity of signals which emerge from the absorbing region.

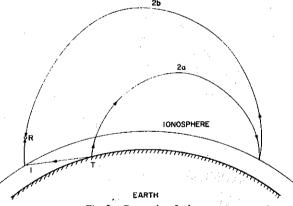


Fig. 3. Ray paths of echoes.

GROUP DELAYS

By noting the delays of the signals, it was usually possible to deduce the general nature of the path of propagation, and to give a picture of the ionosphere consistent with other observations. The "direct" signals (least delay, such as path 1 in Fig. 3) in the northern hemisphere are the

simplest signals to interpret. A typical mid-day delay is 85 m at a height of 600 km. If we assume a Chapman distribution with a maximum plasma frequency of 12 Mc/s at height of 300 km, we obtain a scale height of roughly 120 km. This value is not inconsistent with the average value of 100 km estimated from a number of rocket measurements.⁵

The southern hemisphere group delay measurements show both a direct signal and a "one-hop" whistler mode signal (such as path 2a in Fig. 3). Figure 4 shows the delay data as a function of local mean time. Also shown are the average values of f_0F2 at the equator. The

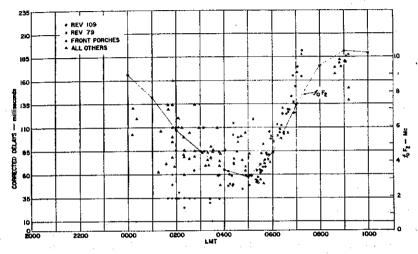


Fig. 4. Group delay and foF2 vs. LMT.

correlation is reasonably good especially in the sunrise period. The data indicate that the scale height increases with time from 80 km at 5.30 a.m. to 100 km at 7 a.m. An interesting result of the analysis is that the exact ray path need not be known to determine the scale height. If the ray path is longer it must reach a higher altitude but the reduced density at the higher altitudes gives less delay per unit path length. These factors almost completely cancel one another at the altitudes involved.

Measurements of the Doppler frequency shift of these signals show a scatter of a few cycles per second with zero mean until about local sunrise. At this time the average frequency suddenly decreases by as much as 7 c/s, where it remains until signals are lost. The observed frequency shift is consistent with the value predicted from the rate of change of group delay using the quasi-longitudinal approximation of the magnetoionic theory.

Strong two-hop whistler mode signals were frequently observed in the northern hemisphere (Fig. 5). The two ray paths which are postulated to explain the signals are shown in Fig. 3. From Figs. 2 and 3 it is seen

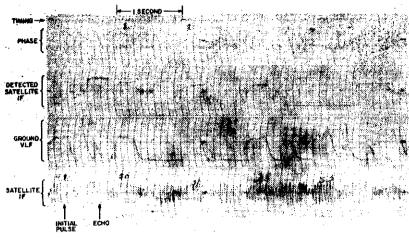


Fig. 5. Echoes on revolution 513.

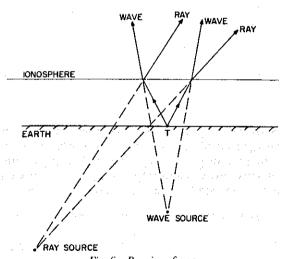


Fig. 6. Beaming of rays.

that the direct signal penetrates the ionosphere with an unfavourable wave-normal angle (see the previous section), and therefore suffers relatively more attenuation. To explain the relatively high strength of the echo, it is necessary to postulate that the first segment of the two-hop path begins close by but to the south of the transmitter. Furthermore, as indicated in Fig. 6, the high refractive index and the strong

anisotropy of the ionosphere in the whistler mode combine to restrict markedly the divergence of the rays. Using a reasonable model of the ionosphere, the calculated group delay is in good agreement with that measured.

CONCLUSIONS

- (1) The LOFTI-I data clearly show that ducts of enhanced ionization are not required for the propagation of whistler-mode signals to a satellite.
- (2) The differential attenuation measurement demonstrates the pronounced effect of the tilt of the earth's magnetic field on attenuation rate. The results are qualitatively in agreement with the predictions of magnetoionic theory. Note that little or no effect would be expected with either a vertical or a horizontal magnetic field.
- (3) Two-hop whistler mode propagation to a satellite yields signals that are often comparable in amplitude to the direct signal. The echo strength may even, on occasion, exceed the direct signal in amplitude.
- (4) Analysis of the group delay data shows that a scale factor on electron distribution can be obtained from this type of experiment.

REFERENCES

- ¹ LEIPHART, J. P., ZEEK, R. W., BEARCE, L. S., and TOTH, E., Penetration of the ionosphere by very-low-frequency radio signals—interim results of the LOFTI I experiment, *Proc. Inst. Radio Engrs.*, N.Y. 50, 6-17, 1962.
- ² STOREY, L. R. O., An investigation of whistling atmospherics, *Phil. Trans. A*, **246**, 113-141, 1953.
- ³ YABROFF, I. W., Computation of whistler ray paths, J. Res. Nat. Bur. Stand. 65D, 5, 485-505 1961.
- 4 Budden, K. G., Radio Waves in Ionosphere, Cambridge University Press, 1961.
- ⁶ WRIGHT, J. W., A model of the F-region above h_{max} F2, J. Geophys. Res. 65, 1, 185-191, 1960.